Limit on Lorentz and CPT Violation of the Neutron Using a Two-Species Noble-Gas Maser

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A search for sidereal variations in the frequency difference between co-located 129 Xe and 3 He Zeeman masers sets the most stringent limit to date on leading-order Lorentz and CPT violation involving the neutron, consistent with no effect at the level of 10^{-31} GeV.

Lorentz symmetry is a fundamental feature of modern descriptions of nature, including both the standard model of particle physics and general relativity. However, these realistic theories are believed to be the low-energy limit of a single fundamental theory at the Planck scale. Even if the underlying theory is Lorentz invariant, spontaneous symmetry breaking might result in small apparent violations of Lorentz invariance at an observable level. Experimental investigations of the validity of Lorentz symmetry therefore provide valuable tests of the framework of modern theoretical physics.

Clock-comparison experiments [1–6] serve as sensitive probes of rotation invariance and hence of Lorentz symmetry, essentially by bounding the frequency variation of a clock as its orientation changes. In practice, the most precise limits are obtained by comparing the frequencies of two different co-located clocks as they rotate with the Earth. Typically, the clocks are electromagnetic signals emitted or absorbed on hyperfine or Zeeman transitions.

Here, we report on a search for sidereal variations in the frequency of co-located 129 Xe and 3 He masers, both operating on nuclear spin-1/2 Zeeman transitions. In the context of a general standard-model extension allowing for the possibility of Lorentz and CPT violation [7,8], the 129 Xe/ 3 He-maser experiment sets the most stringent limit to date on leading order Lorentz and CPT violation of the neutron: about 10^{-31} GeV, or more than six times better than the best previous measurements [9].

The standard-model extension motivating this experiment emerges from any underlying theory that generates the standard model and contains spontaneous Lorentz violation [10]. For example, this might occur in string theory [11]. The standard-model extension maintains theoretically desirable properties of the usual standard model [8]. Its formulation at the level of the known elementary particles is a key feature enabling quantitative comparison of a wide array of tests of Lorentz and CPT symmetry. In this context, theoretical studies have been performed to investigate the sensitivity of clock-comparison experiments [9], tests of QED in Penning traps [12], experiments with a spin-polarized torsion pendulum [13], hydrogen-antihydrogen spectroscopy [14], studies of photon birefringence in the vacuum [15,8,16], observations of

muons [17], measurements of neutral-meson oscillations [7,18], studies of the baryon asymmetry [19], and observations of cosmic rays [20].

In the context of the standard-model extension, the most sensitive prior clock-comparison experiment is the ¹⁹⁹Hg/¹³³Cs comparison of Hunter, Lamoreaux et al. Recent experimental work motivated by the standard-model extension includes Penning trap tests by Gabrielse et al. on the antiproton and the H⁻ ion [21] and by Dehmelt et al. on the electron and positron [22,23]. A reanalysis by Adelberger, Gundlach, Heckel, and co-workers of existing data from a spin-polarized torsion pendulum experiment [24,25] sets the most stringent bound to date on Lorentz and CPT violation of the electron, at about 10^{-28} GeV [26]. A recent Lorentzsymmetry test using hydrogen masers has searched for Zeeman-frequency sidereal variations, placing a bound on Lorentz violation at the level of 10^{-27} GeV [27]. Together with the results of Ref. [26], this implies an improved clean limit of 10^{-27} GeV on Lorentz-violating couplings involving the proton. Also, the KTeV experiment at Fermilab and the OPAL and DELPHI collaborations at CERN have constrained possible Lorentz- and CPTviolating effects in the K and B_d systems [28,29].

The standard-model extension predicts that the leading-order Lorentz- and CPT-violating correction to the 3 He-maser frequency, using the 129 Xe maser as a comagnetometer, is [9]:

$$2\pi \left| \delta \nu_J \right| = \left| -3.5 \tilde{b}_J^n + 0.012 \tilde{d}_J^n + 0.012 \tilde{g}_J^n \right|. \tag{1}$$

Here, \tilde{b}_J^n , \tilde{d}_J^n , and \tilde{g}_J^n are small parameters characterizing the strength of Lorentz-violating couplings of the neutron to possible background tensor fields that may arise from spontaneous symmetry breaking in a fundamental theory. The couplings associated with \tilde{b}_J^n and \tilde{g}_J^n also violate CPT. All three parameters are linear combinations of more basic quantities in the underlying relativistic Lagrangian of the standard-model extension [9].

In the analysis leading to Eq. (1), the Lorentz-violating coupling of either the ${}^{3}\text{He}$ or ${}^{129}\text{Xe}$ nucleus has been taken as that of a single ${}^{1}S_{1/2}$ valence neutron in a Schmidt model [9]. The parameters appearing in Eq.

(1) are therefore associated only with neutron couplings, as indicated by the superscript n on \tilde{b}_J^n , \tilde{d}_J^n , \tilde{g}_J^n . Equation (1) also assumes that the applied magnetic field, which sets the quantization axis of the experiment, is directed east-west in the Earth's reference frame. The subscript J = X, Y indicates components in the sidereal reference frame that are orthogonal to the Earth's axis of rotation.

The design and operation of the two-species ¹²⁹Xe/³He maser has been discussed in recent publications [30,31]. Here, we give a brief review. The two-species maser contains dense co-located ensembles of ³He and ¹²⁹Xe atoms. Each ensemble performs an active maser oscillation on its nuclear spin-1/2 Zeeman transition at approximately 4.9 kHz for ³He and 1.7 kHz for ¹²⁹Xe in a static magnetic field of 1.5 gauss. This two-species maser operation can be maintained indefinitely. The population inversions for the two maser ensembles are created by spinexchange collisions between the noble-gas atoms and optically pumped Rb vapor [32,33]. The ¹²⁹Xe/³He maser has two glass chambers, one acting as the spin exchange "pump bulb" and the other serving as the "maser bulb." This two chamber configuration permits the combination of physical conditions necessary for a high flux of spinpolarized noble gas atoms into the maser bulb, while also maintaining ³He- and ¹²⁹Xe-maser oscillations with good frequency stability: stability of about 100 nHz is typical for measurement intervals larger than about an hour [31]. Either noble-gas species can serve as a precision magnetometer to stabilize the system's static magnetic field, while the other species is employed as a sensitive probe for Lorentz- and CPT-violating interactions.

We used the ¹²⁹Xe/³He maser to search for a Lorentz-violation signature by monitoring the relative phases and Larmor frequencies of the co-located ³He and ¹²⁹Xe masers as the laboratory reference frame rotated with respect to the fixed stars. The system was operated with the ¹²⁹Xe maser as the co-magnetometer, the ³He maser free-running, and the quantization axis directed eastwest in the Earth's reference frame. To leading order, the standard-model extension predicts that the Lorentz-violating frequency shifts for the ³He and ¹²⁹Xe maser are the same size and sign. Hence, the resultant side-real variation of the ³He maser frequency observed in the laboratory frame takes the form

$$\delta\nu_{He} = \delta\nu_X \cos(\Omega_s t) + \delta\nu_Y \sin(\Omega_s t), \tag{2}$$

where Ω_s is the angular frequency of the sidereal day [34], and the parameters $\delta\nu_J$ given by Eq. (1) represent the net effect of the Lorentz- and CPT-violating couplings on the ³He maser frequency with the ¹²⁹Xe maser acting as a co-magnetometer. The time t was measured in seconds from the beginning of the sidereal day in Cambridge, Massachusetts (longitude -71.11°).

Data collection and analysis were performed as follows. The $^{129}\mathrm{Xe}$ - and $^{3}\mathrm{He}$ -maser signals from an inductive pickup coil were buffered, amplified, and sent to a

pair of digital lock-in detectors. Typical raw-signal levels were about 3 to 5 μ V. All reference signals used in the experiment were derived from the same hydrogenmaser clock, thus eliminating concerns about unmeasurable electronic phase shifts between the reference oscillators. The hydrogen maser operated on the standard hyperfine clock transition, and thus had no leading-order sensitivity to Lorentz and CPT violation [9,14]. Active feedback to the solenoid's magnetic field locked the phase of the ¹²⁹Xe maser to that of a 1.7 kHz reference signal, and thereby isolated the experiment from common-mode systematic effects (such as stray magnetic field fluctuations) that would otherwise shift the frequencies of the noble-gas masers in proportion to the ratio of their magnetic moments. When phase locked, the short- and longterm frequency stability [31] of the ¹²⁹Xe maser was several orders of magnitude better than that of the freerunning ³He maser, so the ¹²⁹Xe Zeeman frequency was treated as constant in the data analysis.

The phase and amplitude of both maser signals were recorded at four-second intervals by the lock-in amplifiers and downloaded for analysis every 23.93 hours. A one-sidereal day run thus contained approximately 21,540 evenly spaced measurements of the relative phases of the two masers. The values of the two coefficients $\delta\nu_X, \delta\nu_Y$ were computed, providing a measure of potential Lorentz-violation for that day's run. Seven additional diagnostic signals were recorded, including the temperatures of the pump bulb, maser bulb, and external resonator; an optical monitor of the Rb magnetization in the pump bulb; the broadband power emitted by the opticalpumping laser-diode array; the ambient room temperature; and the east-west component of the ambient magnetic field. Control loops stabilized the system temperatures to about 10 mK. Two additional control loops were used, feeding back to the optical-pumping laser to reduce systematic effects arising from variations in the density and polarization of Rb in the pump bulb [35].

Small noble-gas polarization-induced frequency shifts were the dominant source of instability (i.e., phase drift) in the free-running ³He maser. For a typical one-day run, the linear-correlation coefficient between ³He phase data and the integrated amplitude of either maser was in the range 0.95 - 0.99. We admitted terms to our data-analysis model to account for this polarization-induced phase drift. The effect of potential Lorentz-violating couplings on the evolution of the ³He phase was expressed in terms of the coefficients $\delta\nu_J$ via integration of Eq. (2), and initial reduction of each one-day run was performed using the minimal fit model

$$\delta\phi_{He} = \phi_0 + 2\pi\nu_0 t + 2\pi\Omega_s^{-1} [\delta\nu_X \sin(\Omega_s t) - \delta\nu_Y \cos(\Omega_s t)], \qquad (3)$$

where the coefficients ϕ_0 and ν_0 account for absolute phase and frequency offsets between the ³He maser and

the ultra-stable reference oscillator. The reduced χ^2 statistic for this fit model was determined, and then additional terms corresponding to quadratic and maser amplitude-correlated phase drift were incorporated into the model if they significantly improved the reduced χ^2 [36]. The coefficients $\delta \nu_X$ and $\delta \nu_Y$ for each one-day run were extracted using a linear least-squares routine on the best-fitting model for that day, which contained at most seven free parameters and thus at least (21500 - 7) degrees of freedom. See Fig. 1 for an example of the residuals from one day's data. As a final check, a faux Lorentzviolating effect of known phase and amplitude was added to the raw data and the analysis was repeated. Data reduction for a given sidereal day was considered successful if the synthetic physics was recovered and there was no change in the covariance matrix generated by the fitting routine.

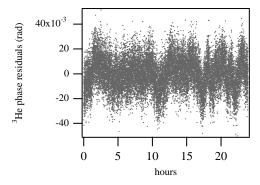


FIG. 1. Typical residuals for the ³He phase data from one sidereal day, calculated using the fit model given in Eq. (3).

Data for this experiment were acquired with three different maser cells over a period of 30 days in April 1999 (cell S3), 24 days in September 1999 (cell E9), and 60 days in February-May 2000 (cell SE3, runs 1 and 2). A total of 90 usable sidereal-day values of $\delta\nu_X$, $\delta\nu_Y$ were obtained. The main magnetic field of the apparatus was reversed about every 10 days to help distinguish possible Lorentz-violating effects from diurnal systematic variations. Field reversal and subsequent re-equilibration of the masers required approximately 24 hours.

Systematic effects resulting from possible diurnally varying ambient magnetic fields would not average away with field reversals. Thus, the effectiveness of the $^{129}\mathrm{Xe}$ co-magnetometer at eliminating such effects was carefully assessed before beginning data acquisition. Since the two maser ensembles do not have perfect spatial overlap, penetration of external magnetic fields through the nested magnetic shields and into the interaction region could induce small frequency shifts in the freerunning $^3\mathrm{He}$ maser despite the presence of the $^{129}\mathrm{Xe}$ comagnetometer. Large coils (~ 2.4 m diameter) surrounding the $^{129}\mathrm{Xe}/^3\mathrm{He}\text{-maser}$ apparatus were used to switch on and off 0.5 G external magnetic fields in the north-

south and east-west directions. A bound on the ratio $|\delta\nu_{He}/\delta B_{external}|$ was obtained. The drifts in the ambient magnetic field near the apparatus were measured to be about 0.2 mG over a typical 24-hour period, resulting in a worst-case shift on the free-running ³He maser of less than 8 nHz, well below the present sensitivity of the experiment to Lorentz and CPT violation. It should be noted that the relative phase between the solar and sidereal day evolved about 2π radians over the course of the experiment (April 1999 to May 2000). Hence, diurnal systematic effects from any source would tend to be reduced by averaging results from all measurement sets.

The measured values of $\delta\nu_J$ exhibited a small dependence on the direction of the solenoidal magnetic field in the laboratory frame. This dependence was most likely due to hysteresis and asymmetry in magnetic interactions between the solenoid and the nested μ -metal shields under field reversal. For each cell, the data for the east and west magnetic-field orientations were analyzed separately to determine mean values and standard errors for $\delta\nu_J$, yielding the results in Table 1. As an example, Fig. 2 shows the single-day values of $\delta\nu_X$ obtained in the first run with cell SE3 in the east field orientation (SE3 E1).

| Cell | $\delta \nu_X ({\rm nHz})$ | $\delta \nu_Y ({\rm nHz})$ |
|--------|-----------------------------|-----------------------------|
| S3 E | 95 ± 118 | 197 ± 114 |
| S3 W | -43 ± 138 | 88 ± 148 |
| E9 E | -86 ± 234 | -194 ± 207 |
| E9 W | -206 ± 186 | -60 ± 134 |
| SE3 E1 | 100 ± 148 | 9 ± 141 |
| SE3~W1 | -1 ± 88 | 62 ± 109 |
| SE3 E2 | -2 ± 180 | 68 ± 107 |
| SE3~W2 | -35 ± 118 | 197 ± 120 |

Table 1: Means and standard errors for $\delta\nu_X$ and $\delta\nu_Y$. Results are displayed for each of the three cells (S3, E9, and SE3) with both east (E) and west (W) orientations of the magnetic field. Two runs were performed for cell SE3.

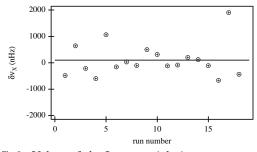


FIG. 2. Values of the Lorentz-violating parameter $\delta\nu_X$ obtained with cell SE3 in the E1 orientation. The horizontal line indicates the mean value for that data set.

The total weighted means and uncertainties for $\delta\nu_X$ and $\delta\nu_Y$ were then formed from all data sets. Finally, the results were used to extract the measured value of $R \equiv \sqrt{\delta\nu_X^2 + \delta\nu_Y^2}$, giving 53 ± 45 nHz (1- σ level).

The size of the coefficients in Eq. (1) indicates that

the $^{129}\mathrm{Xe}/^3\mathrm{He}\text{-}\mathrm{maser}$ experiment is most sensitive to the Lorentz- and CPT-violating coupling associated with $b_J^{\tilde{n}}$. Under the assumption of negligible \tilde{d}_J^n and \tilde{g}_J^n [9], the above experimental result for R corresponds to a value for $\tilde{b}_\perp^n \equiv \sqrt{(\tilde{b}_X^n)^2 + (\tilde{b}_Y^n)^2} = (4.0 \pm 3.3) \times 10^{-31} \, \mathrm{GeV}$. This result is consistent with no Lorentz- and CPT-violating effect, given reasonable assumptions about the probability distribution for R [37]. It represents the most stringent limit to date on possible Lorentz- and CPT-violating couplings involving the neutron and is more than six times better than the best previous measurements [9].

We are planning improved Lorentz and CPT tests using noble-gas masers. Upgrading laser and temperature control and acquiring a larger data set could better the present Lorentz and CPT constraint from the ¹²⁹Xe/³He system by up to an order of magnitude. Also, a new two-species Zeeman maser using ³He and ²¹Ne might provide even greater improvements to constraints on neutron parameters [9,38].

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